

RDMthodfor Unsteady Linear Advection Problem

Consider the advection equation for linear PDE:

$$u_t + \vec{a} \cdot \nabla u = 0$$

This equation can be discretised for the numerical calculation using RD method as following:

$$\sum_{T \in \cup \Delta_i} \left(\sum_{j \in T} m_{ij}^T \frac{u_i^{n+1} - u_i^n}{\Delta t} + \beta_i \phi^T \right) = 0$$

spatial / steady residual

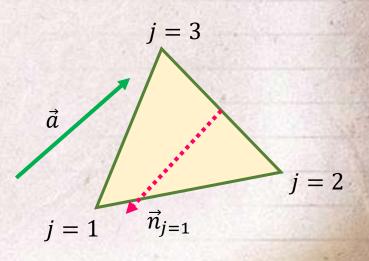
unsteady residual

The spatial residual is computed as

$$\phi^T = \sum_{j \in T} k_j u_j$$

where the inflow parameter is

$$k_j = \frac{1}{2}\vec{a}\cdot\vec{n}_j$$





General Equations of Euler's System

$$\vec{U}_t + \vec{F}_x(\vec{U}) + \vec{G}_y(\vec{U}) = 0$$

where \vec{U} is the primitive variables and the fluxes $F(\vec{U})$ and $G(\vec{U})$ are

$$\vec{U} = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ E \end{pmatrix} \qquad \vec{F}(\vec{U}) = \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ u(E+p) \end{pmatrix} \qquad \vec{G}(\vec{U}) = \begin{pmatrix} \rho v \\ \rho uv \\ \rho v^2 + p \\ v(E+p) \end{pmatrix}$$

where p is the pressure

$$p = (\gamma - 1) \left[E - \frac{1}{2} \rho (u^2 + v^2) \right]$$

Euler's System is in a more general form called the conservation law.

Eq. (1): conservation of mass

Eq. (2) : conservation of momentum in x-direction

Eq. (3) : conservation of momentum in y-direction

Eq. (4): conservation of energy

In fact, advection problem is just a specific case for conservation law.

Let us link the linear advection problem to Euler's system:

Euler's system

$$\vec{U}_t + \vec{F}_x(\vec{U}) + \vec{G}_y(\vec{U}) = 0$$

$$\vec{U}_t + \frac{\partial \vec{F}(\vec{U})}{\partial \vec{U}} \frac{\partial \vec{U}}{\partial x} + \frac{\partial \vec{G}(\vec{U})}{\partial \vec{U}} \frac{\partial \vec{U}}{\partial y} = 0$$

Linear Advection

$$u_t + \vec{a} \cdot \nabla u = 0$$

$$u_t + a_x \frac{\partial u}{\partial x} + a_y \frac{\partial u}{\partial y} = 0$$

The analogy of the "fluxes' speed" in Euler's system is called the Jacobian.



The Jacobian matrix can be found as following:

$$\vec{U} = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ E \end{pmatrix} = \begin{pmatrix} U_1 \\ U_2 \\ U_3 \\ U_4 \end{pmatrix}$$

$$\vec{F}(\vec{U}) = \begin{pmatrix} f_1(\vec{U}) \\ f_2(\vec{U}) \\ f_3(\vec{U}) \\ f_4(\vec{U}) \end{pmatrix} \qquad \vec{G}(\vec{U}) = \begin{pmatrix} g_1(\vec{U}) \\ g_2(\vec{U}) \\ g_3(\vec{U}) \\ g_4(\vec{U}) \end{pmatrix}$$

$$\frac{\partial \vec{f}_{1}(\vec{U})}{\partial \vec{U}} = \begin{pmatrix}
\frac{\partial f_{1}(\vec{U})}{\partial U_{1}} & \frac{\partial f_{1}(\vec{U})}{\partial U_{2}} & \frac{\partial f_{1}(\vec{U})}{\partial U_{3}} & \frac{\partial f_{1}(\vec{U})}{\partial U_{4}} \\
\frac{\partial f_{2}(\vec{U})}{\partial U_{1}} & \frac{\partial f_{2}(\vec{U})}{\partial U_{2}} & \frac{\partial f_{2}(\vec{U})}{\partial U_{3}} & \frac{\partial f_{2}(\vec{U})}{\partial U_{4}} \\
\frac{\partial f_{3}(\vec{U})}{\partial U_{1}} & \frac{\partial f_{3}(\vec{U})}{\partial U_{2}} & \frac{\partial f_{3}(\vec{U})}{\partial U_{3}} & \frac{\partial f_{3}(\vec{U})}{\partial U_{4}} \\
\frac{\partial f_{4}(\vec{U})}{\partial U_{1}} & \frac{\partial f_{4}(\vec{U})}{\partial U_{2}} & \frac{\partial f_{4}(\vec{U})}{\partial U_{3}} & \frac{\partial f_{4}(\vec{U})}{\partial U_{4}}
\end{pmatrix}$$

Similar procedure applies for $\frac{\partial \vec{G}(\vec{U})}{\partial \vec{U}}$

InflowNtrixk; for Elle's Stem

Euler's system

Linear Advection

inflow matrix:
$$K_j = \frac{1}{2} A |\vec{n}_j|$$

inflow parameter:
$$k_j = \frac{1}{2} \vec{a} \cdot \vec{n}_j$$

Jacobian:
$$A = \frac{\partial \vec{F}(\vec{U})}{\partial \vec{U}} n_x + \frac{\partial \vec{G}(\vec{U})}{\partial \vec{U}} n_y$$

Speed:
$$\vec{a} = a_x \hat{\imath} + a_y \hat{\jmath}$$

The Jacobian matrix is given as

$$A = \begin{pmatrix} 0 & n_{x} & n_{y} & 0 \\ \frac{1}{2}(\gamma - 1)q^{2}n_{x} - u\vec{u} \cdot \vec{n} & -(\gamma - 2)un_{x} + \vec{u} \cdot \vec{n} & -(\gamma - 1)vn_{x} + un_{y} & (\gamma - 1)n_{x} \\ \frac{1}{2}(\gamma - 1)q^{2}n_{y} - v\vec{u} \cdot \vec{n} & -(\gamma - 1)un_{y} + vn_{x} & -(\gamma - 2)vn_{y} + \vec{u} \cdot \vec{n} & (\gamma - 1)n_{y} \\ \vec{u} \cdot \vec{n} \begin{bmatrix} \frac{1}{2}(\gamma - 1)q^{2} - H \end{bmatrix} & Hn_{x} - (\gamma - 1)u\vec{u} \cdot \vec{n} & Hn_{y} - (\gamma - 1)v\vec{u} \cdot \vec{n} & \gamma \vec{u} \cdot \vec{n} \end{pmatrix}$$



where H is the enthalpy, $H = \frac{p + E}{\rho}$

$$q^2 = u^2 + v^2$$

Dagmaistility of the Jackian Matrix

Euler's system

Linear Advection

inflow matrix:
$$K_j = \frac{1}{2}A |\vec{n}_j|$$

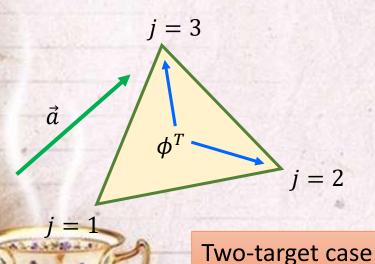
$$K_j = \frac{1}{2} A \left| \vec{n}_j \right|$$

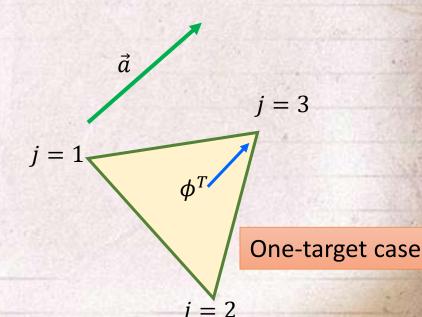
inflow parameter:
$$k_j = \frac{1}{2} \vec{a} \cdot \vec{n}_j$$

For the LDA scheme, the distribution coefficient is given as:

$$\beta_i = \frac{k_i^+}{\sum_{j \in T} k_j^+}$$

$$k_i^+ = \max(0, k_j)$$





The question now is how to find the corresponding positive and negative inflow matrix for Euler's system?

$$K^+$$
 ; K^-

Euler's system is a hyperbolic PDE because its Jacobian matrix is diagonalisable

A matrix is said to be diagonalisable iff

where D is a diagonal matrix

$$B = C^{-1}DC$$

$$\downarrow \qquad \qquad \downarrow \qquad \qquad \downarrow$$

$$A = RAR^{-1}$$

A is the Jacobian, R is the right eigenvectors,

 Λ is an diagonal matrix which its diagonal entries are the eigenvalues

The Jacobian matrix can be expressed as an eigenvalue problem:

$$AR^{(i)} = \lambda_i R^{(i)}$$





There are 4 eigenvalues and therefore 4 eigenvectors, arranged in increasing magnitude of the eigenvalues give

$$\Lambda = \begin{pmatrix} \lambda_1 & 0 & 0 & 0 \\ 0 & \lambda_2 & 0 & 0 \\ 0 & 0 & \lambda_3 & 0 \\ 0 & 0 & 0 & \lambda_4 \end{pmatrix}$$

$$R = (\vec{R}^{(1)}, \vec{R}^{(2)}, \vec{R}^{(3)}, \vec{R}^{(4)})$$

where

$$\lambda_1 = \vec{u} \cdot \vec{n} - c$$

$$\lambda_1 = \vec{u} \cdot \vec{n} - c$$
 $\lambda_2 = \lambda_3 = \vec{u} \cdot \vec{n}$

$$\lambda_4 = \vec{u} \cdot \vec{n} + c$$

$$R = \begin{pmatrix} 1 & 0 & 1 & 1 \\ u - cn_x & -n_y & u & u + cn_x \\ v - cn_y & n_x & v & v + cn_y \\ H - c\vec{u} \cdot \vec{n} & -un_y + vn_x & \frac{u^2 + v^2}{2} & H + c\vec{u} \cdot \vec{n} \end{pmatrix}$$

where c is the sound speed and is given by : $c^2 = (\gamma - 1) \left[H - \frac{1}{2} q^2 \right]$

$$c^2 = (\gamma - 1) \left[H - \frac{1}{2} q^2 \right]$$



 R^{-1} is the inverse of the right eigenvectors or just simply known as the left eigenvectors.

Euler's system

Linear Advection

inflow matrix:
$$K_j = \frac{1}{2} A |\vec{n}_j|$$

inflow parameter:
$$k_j = \frac{1}{2} \vec{a} \cdot \vec{n}_j$$

$$A^+ = R\Lambda^+R^{-1}$$
$$A^- = R\Lambda^-R^{-1}$$

$$k_j^+ = \begin{cases} k_j & if & \overline{\vec{a}} \cdot \vec{n}_j > 0 \\ 0 & if & \overline{\vec{a}} \cdot \vec{n}_j \le 0 \end{cases}$$

$$k_j^- = \begin{cases} k_j & if & \overline{\vec{a}} \cdot \vec{n}_j < 0\\ 0 & if & \overline{\vec{a}} \cdot \vec{n}_j \ge 0 \end{cases}$$

The distribution matrix then becomes:

$$B_i = K_i^+ \left(\sum_{j \in T} K_j^+\right)^{-1}$$

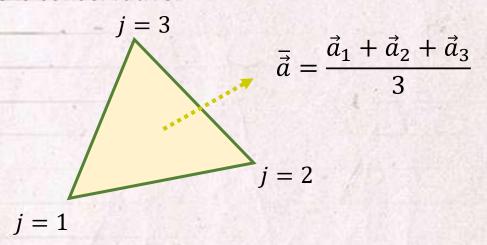
and the total spatial residual of an element is:



$$\Phi^{T} = \iint_{T} [\vec{F}_{x}(\vec{U}) + \vec{G}_{x}(\vec{U})] dxdy = \iint_{T} [\vec{F}(\vec{U})dy + \vec{G}(\vec{U})dx]$$
$$= \frac{1}{2} \sum_{j \in T} [\vec{F}(\vec{U}_{j}) + \vec{G}(\vec{U}_{j})] \cdot \vec{n}_{j}$$

The Aeraged Jackbian of an Herrent T

In scalar advection problem (both linear advection and Burger's equation), the speed vector has to be assumed constant in each element so that the RD scheme is conservative.



This in turn implies that all the primitive variables of Jacobian A must be assumed to be constant within each element T.





$$\lambda_1 = \bar{\vec{u}} \cdot \vec{n} - \bar{c}$$

$$\lambda_2 = \lambda_3 = \overline{\vec{u}} \cdot \vec{n}$$

$$\lambda_4 = \bar{\vec{u}} \cdot \vec{n} + \bar{c}$$

Right Eigenvectors:

$$R = \begin{pmatrix} 1 & 0 & 1 & 1 \\ \bar{u} - \bar{c}n_{x} & -n_{y} & \bar{u} & \bar{u} + \bar{c}n_{x} \\ \bar{v} - \bar{c}n_{y} & n_{x} & \bar{v} & \bar{v} + \bar{c}n_{y} \\ \bar{H} - \bar{c}\bar{\vec{u}} \cdot \vec{n} & -\bar{u}n_{y} + \bar{v}n_{x} & \frac{\bar{u}^{2} + \bar{v}^{2}}{2} & \bar{H} + \bar{c}\bar{\vec{u}} \cdot \vec{n} \end{pmatrix}$$

First of all, find the averaged value for all the terms contained in the primitive variables:

$$\bar{\rho} = \frac{\rho_1 + \rho_2 + \rho_3}{3}$$

$$\bar{u} = \frac{u_1 + u_2 + u_3}{3}$$

$$\bar{v} = \frac{v_1 + v_2 + v_3}{3}$$

$$\bar{\rho} = \frac{\rho_1 + \rho_2 + \rho_3}{3}$$
 $\bar{u} = \frac{u_1 + u_2 + u_3}{3}$ $\bar{v} = \frac{v_1 + v_2 + v_3}{3}$ $\bar{E} = \frac{E_1 + E_2 + E_3}{3}$

and the use these averaged values to compute for the following variables:

$$\bar{p} = (\gamma - 1) \left[\bar{E} - \frac{1}{2} \bar{\rho} (\bar{u}^2 + \bar{v}^2) \right]$$

$$\overline{H} = \frac{\overline{p} + E}{\overline{o}}$$

$$\bar{q}^2 = \bar{u}^2 + \bar{v}^2$$



$$\bar{c}^2 = (\gamma - 1) \left[\bar{H} - \frac{1}{2} \bar{q}^2 \right]$$

Generalising the Salar Advettion Equation to the Eller's System—the Mass Matrix

Scalar Advection

$$\sum_{T \in \cup \Delta_i} \left(\sum_{j \in T} m_{ij}^T \frac{u_i^{n+1} - u_i^n}{\Delta t} + \beta_i \phi^T \right) = 0$$

$$m_{ij}^T = \frac{S_T}{3} \beta_i (\delta_{ij} + 1 - \beta_j)$$

Euler's system

$$B_i = K_i^+ \left(\sum_{j \in T} K_j^+\right)^{-1}$$

$$\Phi^{T} = \frac{1}{2} \sum_{j \in T} \left[\vec{F}(\vec{U}_j) + \vec{G}(\vec{U}_j) \right] \cdot \vec{n}_j$$

$$\sum_{T \in \cup \Delta_i} \left(\sum_{j \in T} M_{ij}^T \frac{\vec{U}_i^{n+1} - \vec{U}_i^n}{\Delta t} + B_i \Phi^T \right) = 0$$

identity matrix

$$M_{ij}^{T} = \frac{S_{T}}{3} B_{i} \left(\delta_{ij} I + I - B_{j} \right)$$



Remarks and Corrections for last presentation (9/3/2016):

The semi-discrete model for scalar advection problem reads,

$$\left[\frac{u_i^{n+1} - u_i^n}{\Delta t} + \mathcal{O}(\Delta t^2) \right]^{n + \frac{1}{2}} + \left[\vec{a} \cdot \nabla u \right]^{n + \frac{1}{2}} = 0$$

The expansion of the Taylor's series at time level $n + \frac{1}{2}$ is to preserve the second order accuracy in the discretisation of time derivative.

<u>In previous presentation</u>, where the advection speed for Burgers' equation $\vec{a}(u)$ depend on u and it varies with time:

$$\sum_{T \in \cup \Delta_i} \left(\sum_{j \in T} m_{ij}^{T, n + \frac{1}{2}} \frac{u_i^{n+1} - u_i^n}{\Delta t} + \frac{1}{2} \beta_i^{n + \frac{1}{2}} (\phi^T(u^{n+1}) + \phi^T(u^n)) \right) = 0$$



In actual case, implicit scheme with mass matrix should be implemented as follow

$$\sum_{T \in \cup \Delta_i} \left(\sum_{j \in T} \left(m_{ij}^{T,n+1} \frac{u_i^{n+1}}{\Delta t} - m_{ij}^{T,n} \frac{u_i^n}{\Delta t} \right) + \frac{1}{2} \left(\beta_i^{n+1} \phi^T(u^{n+1}) + \beta_i^n \phi^T(u^n) \right) \right) = 0$$

The mass matrix is constructed based on the assumption of linear preserving scheme.

Examples of RD framework for unsteady case using mass matrix are:

Rossiello:

UCV, LDA, monotone FCT procedure

Ricchiuto, Abgrall:

LDA, blended LDA-N, SU, central-blended (with Lax-Friedrich's dissipation)

All these scheme are linear preserving.

Linear preservation means that the distribution coefficient must be bounded.

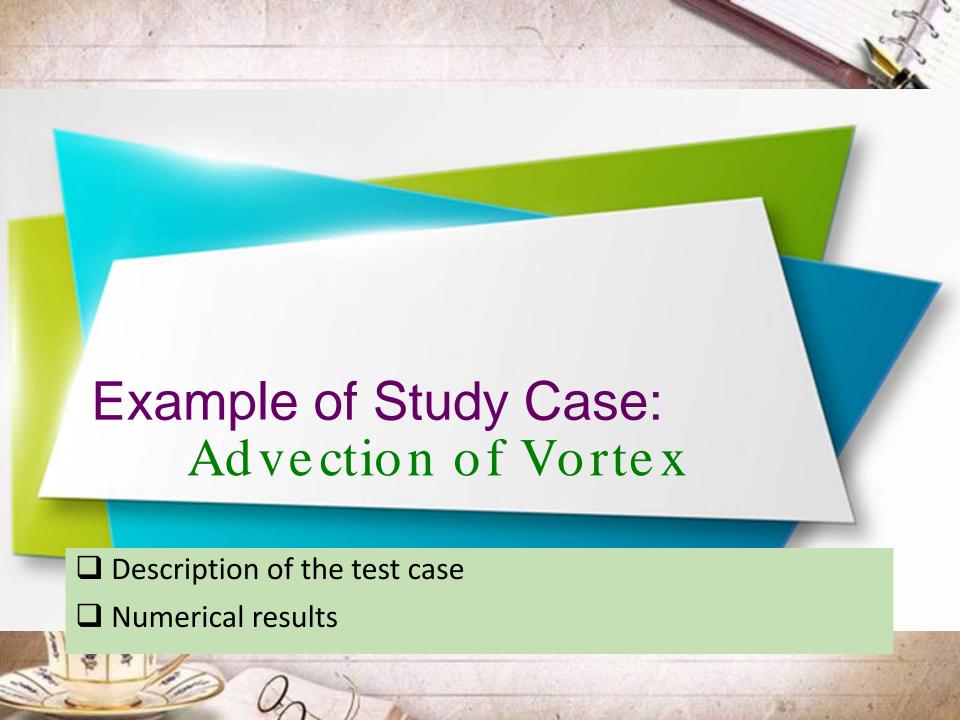
$$0 \le \beta \le 1$$

The linear preserving property of the distribution coefficient will ensure the existence of mass matrix m_{ij}^{T}

Streamline upwind mass matrix

$$m_{ij}^T = \frac{S_T}{3} \begin{bmatrix} \beta_1^T (2 - \beta_1^T) & \beta_1^T (1 - \beta_2^T) & \beta_1^T (1 - \beta_3^T) \\ \beta_2^T (1 - \beta_1^T) & \beta_2^T (2 - \beta_2^T) & \beta_2^T (1 - \beta_3^T) \\ \beta_3^T (1 - \beta_1^T) & \beta_3^T (1 - \beta_2^T) & \beta_3^T (2 - \beta_3^T) \end{bmatrix}$$



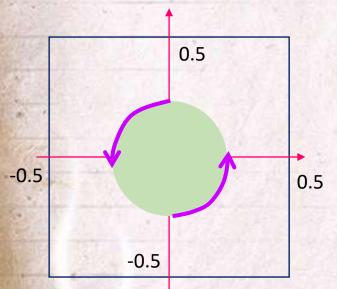


$$\vec{U}_t + \vec{F}_x(\vec{U}) + \vec{G}_y(\vec{U}) = 0$$

$$\vec{U} = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ E \end{pmatrix}$$

$$\vec{U} = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ E \end{pmatrix} \qquad \vec{F}(\vec{U}) = \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho uv \\ u(E+p) \end{pmatrix} \qquad \vec{G}(\vec{U}) = \begin{pmatrix} \rho v \\ \rho uv \\ \rho v^2 + p \\ v(E+p) \end{pmatrix}$$

$$\vec{G}(\vec{U}) = \begin{pmatrix} \rho v \\ \rho u v \\ \rho v^2 + p \\ v(E+p) \end{pmatrix}$$



Density is assumed to be constant throughout the domain

$$\rho = 1.4$$

The centre of vortex is initially set to be

$$(x_c, y_c) = (0.0)$$

$$r = \sqrt{(x - x_c)^2 + (y - y_c)^2}$$

The flow velocity is given as

$$\vec{v} = \vec{v}_m + \vec{v}_p$$

Mean stream velocity: $\vec{v}_m = (6,0)$

Circumferential perturbation:

$$\vec{v}_p = \begin{cases} 15(\cos(4\pi r) + 1)(-(y - y_c), (x - x_c)), & r < 0.25\\ 0, & otherwise \end{cases}$$

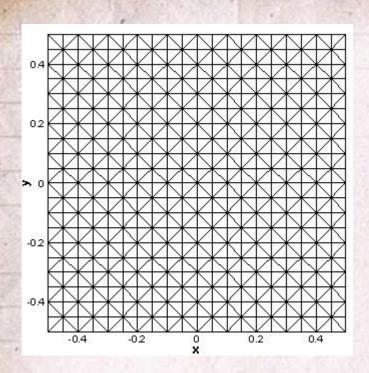
The pressure is given as

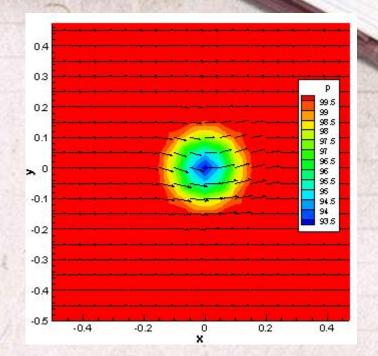
$$p=p_m+p_p$$

$$p_m=100 \qquad p_p= \begin{cases} \Delta p &, r<0.25 \\ 0 &, otherwise \end{cases}$$

$$\Delta p = \frac{15^2 \rho}{(4\pi)^2} \left(2\cos(4\pi r) + 8\pi r \sin(4\pi r) + \frac{\cos(8\pi r)}{8} + \pi r \sin(8\pi r) + 12\pi^2 r^2 \right) + C$$

$$C = -11.02544849$$





Global time step:

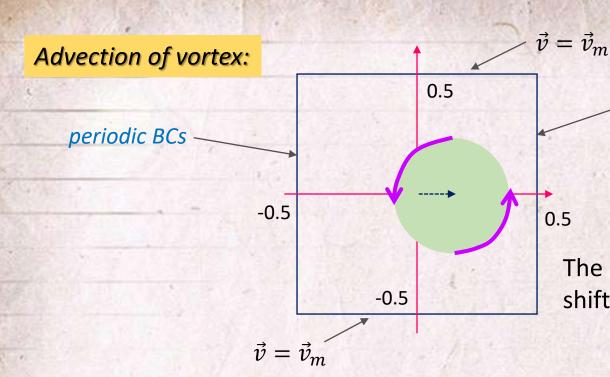
$$\Delta t = \min_{i} \Delta t_i$$

$$\Delta t_i = CFL \frac{S_i}{\sum_{T \in \cup \Delta_i} \alpha^T}$$

$$\alpha^{T} = \frac{1}{2} \max_{j \in T} (\|\vec{v}_{j} + c_{j}\|) |\vec{n}_{j}|$$



$$c_j = \sqrt{\frac{\gamma p_j}{\rho_j}}$$



The centre of vortex has to be shifted as time elapsed:

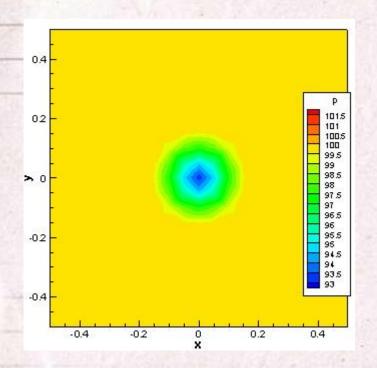
$$(x_c, y_c) = \Delta t \ \vec{v}_m$$

$$\vec{U} = \begin{pmatrix} \rho \\ \rho u \\ \rho v \\ E \end{pmatrix}$$

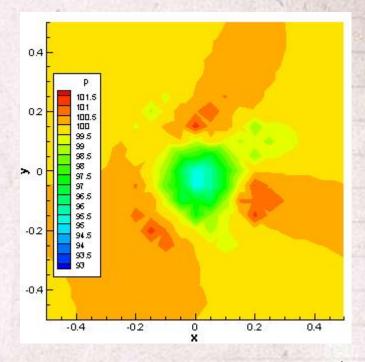
fixed

Unknown, and the pressure is to be calculated from

$$p = (\gamma - 1) \left[E - \frac{1}{2} \rho (u^2 + v^2) \right]$$



Exact pressure at $t = \frac{1}{6}$



Numerical results at $t = \frac{1}{6}$



an)